

# Line-by-line spectroscopic parameters of HFC-32 ro-vibrational transitions within the atmospheric window around 8.2 $\mu\text{m}$

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## ABSTRACT

Diffuoromethane ( $\text{CH}_2\text{F}_2$ , HFC-32) presents strong ro-vibrational bands within the 8–12  $\mu\text{m}$  atmospheric window and hence it represents a greenhouse gas able of contributing to global warming. Numerous spectroscopic studies have been devoted to this molecule, however, much information on line-by-line parameters, like line intensities and broadening parameters, is still lacking. In this work, line-by-line spectroscopic parameters are retrieved for several  $\text{CH}_2\text{F}_2$  ro-vibrational transitions belonging to the  $\nu_7$  band located around 8.5  $\mu\text{m}$ . Self-broadening as well  $\text{N}_2$ - and  $\text{O}_2$ - broadening experiments are carried out at room temperature by using a tunable diode laser (TDL) spectrometer. The line shape analysis of  $\text{CH}_2\text{F}_2$  self-broadened spectra leads to the determination of resonant frequencies, integrated absorption coefficients and self-broadening parameters, while  $\text{CH}_2\text{F}_2$ - $\text{N}_2$  and  $\text{CH}_2\text{F}_2$ - $\text{O}_2$  broadening coefficients are obtained from foreign-broadening measurements. In addition, the broadening parameters of  $\text{CH}_2\text{F}_2$  in air are derived from the  $\text{N}_2$ - and  $\text{O}_2$ - broadening coefficients. The results of the present work provide fundamental information to measure the concentration profiles of this molecule in the atmosphere through remote sensing spectroscopic techniques.

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## 1. Introduction

Since their introduction chlorofluorocarbons (CFCs) have been employed as propellants, refrigerants, blowing agents and fire extinguisher thanks to their favorable chemical and thermal properties. Nevertheless, CFCs contribute to the destruction of the stratospheric ozone layer and hence, with only a few exceptions, they have been phased out by the Montreal protocol. Among the replacement compounds, hydrofluorocarbons (HFCs) have been considered because of the absence of Cl atoms and the usually lower atmospheric lifetimes than CFCs. However, HFCs contribute to global climate changes since they behave as greenhouse gases. In fact, the C–F group strongly absorbs the infrared radiation within the 8–12  $\mu\text{m}$  atmospheric window, thus altering the Earth's radiative budget. Therefore, the spectroscopic properties of HFCs are the subject of intensive research efforts, aimed at assessing their potential environmental impacts and at providing the required information for their atmospheric detection and modeling (e.g. see Refs. [1–3] and references therein).

For remote sensing spectroscopic applications, the work of laboratory spectroscopy is to provide accurate values of line-by-line

spectroscopic parameters (i.e. transition frequencies, line intensities, pressure broadening and shifting coefficients and eventually, their temperature dependence) and absorption cross sections [4–6], the latter being particularly suitable for species presenting highly congested spectra. Furthermore, pressure broadening and shifting coefficients are strictly related to the intermolecular collisions and the potential driving them and hence their knowledge can provide important information about the chemical-physics of non-reactive scattering events and decay processes in the gas phase (see [7–14] and references therein). Clearly, for atmospheric purposes, the main collisional partners are  $\text{N}_2$  and  $\text{O}_2$ . The line-by-line spectroscopic parameters and the absorption cross sections data of the species of interest are collected into a number of databases that are continuously updated in terms of spectral coverage, accuracy, chemical species and so on.

Among atmospheric trace gas pollutants, difluoromethane ( $\text{CH}_2\text{F}_2$ , HFC-32) is included in the HITRAN database [15], which lists cross-section data. This compound is employed in refrigerant mixtures together with  $\text{CF}_3\text{CH}_3$ ,  $\text{CF}_3\text{CH}_2\text{F}$ , and  $\text{CF}_3\text{CHF}_2$ . Its atmospheric concentration has been determined to be around 3 ppt in 2005 [16], with a growing trend with respect to previous years. Because its environmental relevance,  $\text{CH}_2\text{F}_2$  has been the subject of notable interest from the spectroscopic community (for an overview see [17–19] and references therein). Recently, the  $\text{CH}_2\text{F}_2$

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adsorption over the TiO<sub>2</sub> surface has been studied by coupling diffuse reflectance infrared spectroscopy to periodic calculations rooted into density functional theory, being adsorption the first step involved in gas-phase heterogeneous catalysis aimed at removing atmospheric pollutants [20].

Focusing on absorption cross sections, Smith et al. determined the CH<sub>2</sub>F<sub>2</sub> integrated absorbance cross sections over the 600–1900 cm<sup>-1</sup> spectral range at different temperatures between 203 and 297 K [21], while few years later Beukes and Nicolaisen measured the intensity of the  $\nu_4$  fundamental band at six temperatures within the 212.5–297.0 K interval [22]. Meanwhile, integrated absorption cross sections were obtained with the aim of deriving the radiative forcing and global warming potential [23], and later on, they were re-investigated in the spectral region between 400 and 1600 cm<sup>-1</sup> at 295 K by considering the effect of air broadening on the retrieved values [24]; more recently, the determination of integrated absorption cross sections was extended over the entire medium infrared region, up to 5000 cm<sup>-1</sup> [17].

On the other hand, up to now only one study has dealt with the determination of pressure broadening coefficients of self-, N<sub>2</sub>- and O<sub>2</sub>-broadened CH<sub>2</sub>F<sub>2</sub> [25]. The foreign broadening coefficients of the CH<sub>2</sub>F<sub>2</sub> collisionally perturbed by different damping gases have been employed to derive the dissociation energy of the (CH<sub>2</sub>F<sub>2</sub>)<sub>2</sub> dimer and the experimental results have been benchmarked against predictions from different quantum mechanical methods [26,27].

Since line-by-line spectroscopic parameters assume particular relevance to acquire basic information for the atmospheric modeling of this compound and for remote sensing applications, this work aims at retrieving new data about self-, N<sub>2</sub>-, and O<sub>2</sub>- broadening coefficients of CH<sub>2</sub>F<sub>2</sub> and integrated absorption coefficients. The focus is on lines from the  $\nu_7$  band, which, despite being weaker and less suited for satellite remote sensing than those from the  $\nu_9$  band, could offer useful information for atmospheric applications.

## 2. Experimental details

High-resolution IR spectra of CH<sub>2</sub>F<sub>2</sub> perturbed by either itself or by N<sub>2</sub> and O<sub>2</sub> collisional partners were recorded by using a tunable diode laser (TDL) spectrometer [28] employing two different commercial lead salt lasers emitting within the atmospheric region and centered at 8.3 and 8.1  $\mu$ m, respectively. All the measurements were carried out at room temperature (295  $\pm$  1 K) with the spectrometer operating in a three beam configuration [29]: the first beam passed through the 68.2 ( $\pm$ 0.2) cm absorption cell containing the CH<sub>2</sub>F<sub>2</sub> sample; the other two beams were respectively sent to a Germanium etalon and a second cell filled with SO<sub>2</sub>, both employed for calibration purposes. The three beams were collected by three different HgCdTe detectors, cryogenically cooled at liquid nitrogen temperature, and up to 1024 independent scans were averaged and digitized by a four-channel digital oscilloscope, resulting in spectra with a signal-to-noise ratio better than 1000. The acquired spectral micro-windows, about 0.3–0.5 cm<sup>-1</sup> wide, were wavenumber calibrated by using the frequency of suitable SO<sub>2</sub> lines from high-resolution FTIR spectra [29], resulting in a calibration accuracy of ca.  $4 \times 10^{-4}$  cm<sup>-1</sup> for unblended spectral lines. The transmittance spectra were obtained from the incident and transmitted radiation intensities according to the Beer-Lambert's law, with the incident intensity retrieved by fitting no absorption regions (i.e. the baseline) to a polynomial function.

For the self-broadening measurements, the sample pressure was varied in the range between 25 and 500 Pa, and each spectral micro-window was recorded up to 10 different pressures. For the foreign broadening measurements, spectra were recorded by using a fixed amount of CH<sub>2</sub>F<sub>2</sub>, around 100 Pa, depending on the

intensities of the spectral lines, and increasing the total pressure up to 40 hPa. Pressure measurements were performed by means of the Alcatel ARD 1003 and 1002 capacitance vacuum gauges with a full scale range of 10 hPa and 100 hPa, respectively, each with quoted manufacturer's full scale accuracy of 0.15% and the pressure was found stable during the spectra recording well within 0.02 hPa and 0.2 hPa for self- and foreign-broadening experiments, respectively. Furthermore, following a well consolidated procedure [25,28,29], after checking single mode emission from the laser, the measurements were performed by removing the monochromator to avoid instrumental distortions. A complete equilibration inside the cell was achieved by adopting a 15 min delay between the filling and the recording of the corresponding spectrum. Gas handling operations were performed by using a stainless steel vacuum line and before each measurement, the cell was evacuated to about 10<sup>-4</sup> Pa by means of a diffusion pump backed by a double stage rotary pump. The CH<sub>2</sub>F<sub>2</sub> sample was purchased by Aldrich with a stated purity of 99.7%, and it was used without any further purification. To retrieve line-by-line spectroscopic parameters, the recorded absorption features were fitted by using the VLSF program operating in the multiline mode [30]. Assignment of the spectral lines was performed on the basis of Ref. [19].

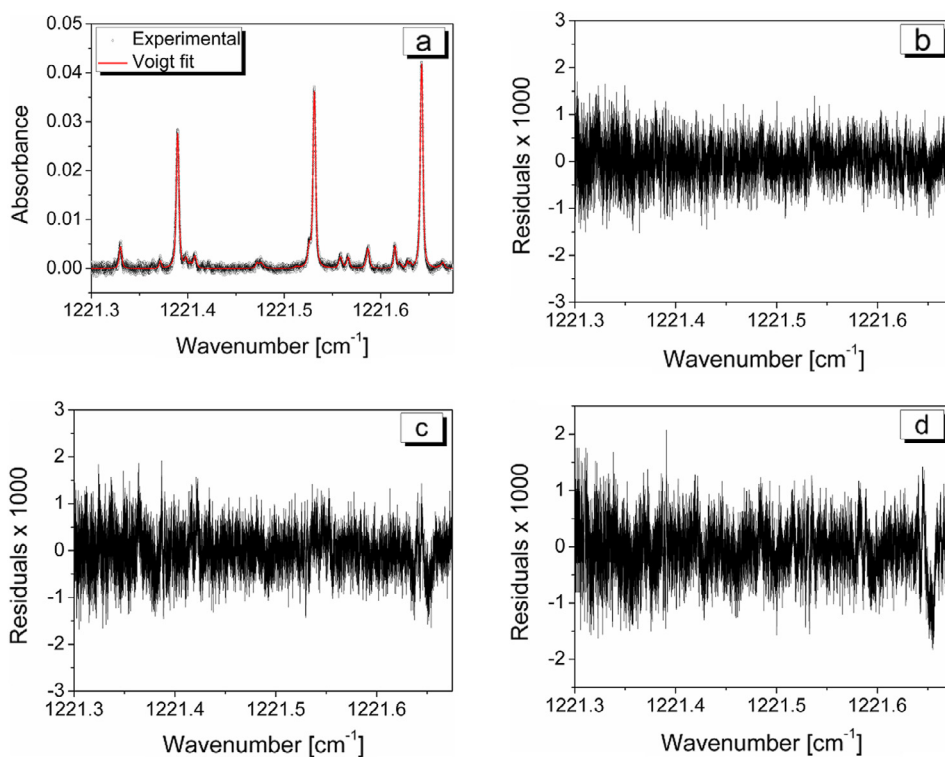
## 3. Results and discussion

The CH<sub>2</sub>F<sub>2</sub> molecule belongs to the C<sub>2v</sub> symmetry point group and it is a near prolate asymmetric rotor as confirmed by the Ray's asymmetry parameter  $\kappa = -0.93$ . It possesses nine normal modes of vibration which can be classified as belonging to A<sub>1</sub> ( $\nu_1$ – $\nu_4$ ), A<sub>2</sub> ( $\nu_5$ ), B<sub>1</sub> ( $\nu_6$ ,  $\nu_7$ ) or B<sub>2</sub> ( $\nu_8$ ,  $\nu_9$ ) symmetry species. The IR spectrum in the 1000–1300 cm<sup>-1</sup> atmospheric spectral region is characterized by the absorptions of the  $\nu_3$ ,  $\nu_5$ ,  $\nu_7$  and  $\nu_9$  fundamental vibrations, which correspond to the CF<sub>2</sub> symmetric stretching, CH<sub>2</sub> twisting, CH<sub>2</sub> rocking and CF<sub>2</sub> asymmetric stretching vibrations, respectively. The corresponding excited vibrational levels form, together with the  $\nu_4 = 2$  level, a resonant pentad interacting through Coriolis and anharmonic resonances [19]. Because of these interactions the  $\nu_5$  vibration, which is expected to be inactive in the IR, gives rise to a weak band centered at 1255.85851 cm<sup>-1</sup>. Among the remaining fundamental bands within the atmospheric window,  $\nu_9$  produces the most intense absorption of the whole CH<sub>2</sub>F<sub>2</sub> IR spectrum centered at 1090.126362 cm<sup>-1</sup>. It is overlapped with the  $\nu_3$  band at 1111.516271 cm<sup>-1</sup>, whose R branch is, in turn severely blended with the absorption of the  $\nu_7$  fundamental resonating at 1178.641412 cm<sup>-1</sup>. These bands are expected to produce transitions in the P, Q and R branches classifiable as even ( $K''_a + K''_c = J''$ ) and odd ( $K''_a + K''_c = J'' + 1$ ) sub-bands.

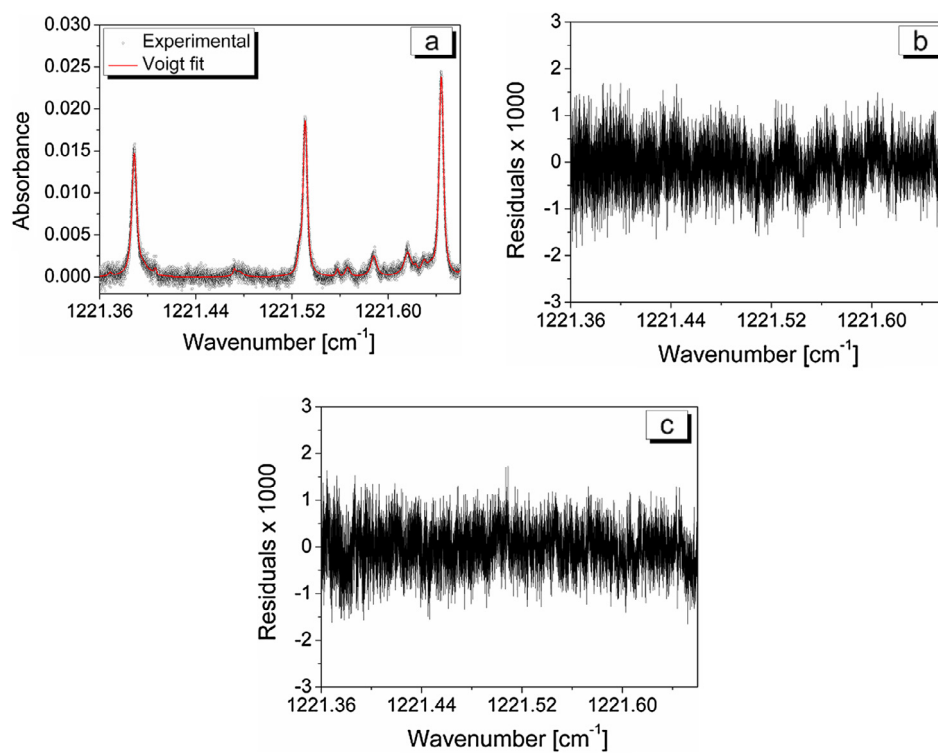
The spectral range covered by the TDL spectrometer matches the ro-vibrational lines stemming from the  $\nu_7$  normal mode and, of these 5, 20 and 35 belong to the P, Q and R branch, respectively. Because the lasers employed for the measurements present a Lorentzian instrumental line shape function, the spectral lines have been fitted to the Voigt profile. During the fits, the Gaussian half widths have been fixed to the molecular contribution  $\gamma_D^{\text{CH}_2\text{F}_2}$  given by

$$\gamma_D^{\text{CH}_2\text{F}_2} = \frac{v_{if}}{c} \left( \frac{2k_B N_a T \ln 2}{m} \right)^{1/2} \quad (1)$$

where  $v_{if}$  is the frequency of the  $f \leftarrow i$  ro-vibrational transition,  $c$  represents the speed of light,  $k_B$  is the Boltzman's constant,  $N_a$  is the Avogadro's constant,  $T$  is the absolute temperature and  $m$  is the mass of the radiating species. An example of the fits performed during self-broadening measurements is given in Fig. 1, showing the spectral region between 1221.30 and 1221.67 cm<sup>-1</sup>, while Fig. 2 presents the CH<sub>2</sub>F<sub>2</sub> spectrum perturbed by similar amounts



**Fig. 1.** (a) High resolution spectrum of self-broadened  $\text{CH}_2\text{F}_2$  between 1221.30 and 1221.67  $\text{cm}^{-1}$  ( $\text{CH}_2\text{F}_2$  pressure = 123.0 Pa; optical path = 68.2 cm;  $T = 295$  K) showing the synthetic trace obtained by employing the Voigt profile. (b) Residuals for  $\text{CH}_2\text{F}_2$  pressure = 46.1 Pa. (c) Residuals for  $\text{CH}_2\text{F}_2$  pressure = 123.0 Pa. (d) Residuals for  $\text{CH}_2\text{F}_2$  pressure = 401.1 Pa.



**Fig. 2.** High-resolution IR spectrum of  $\text{CH}_2\text{F}_2$  between 1221.36 and 1221.66  $\text{cm}^{-1}$  perturbed by similar amounts of  $\text{O}_2$  and  $\text{N}_2$  damping gases. Experimental spectrum and synthetic trace obtained using the Voigt profile ( $\text{CH}_2\text{F}_2$  partial pressure = 99.4 Pa;  $\text{N}_2$  partial pressure = 941 Pa; optical path = 68.2 cm;  $T = 295$  K). (b) Residuals of the fit for  $\text{N}_2$  partial pressure = 941 Pa. (c) Residuals of the fit for  $\text{O}_2$  partial pressure = 881 Pa.

of  $O_2$  and  $N_2$  damping gases, about 900 Pa, in the 1221.36 and 1221.66  $cm^{-1}$  range. These figures illustrate two important aspects concerning the line shape analysis performed on the high resolution spectra of  $CH_2F_2$ . The first one is the presence of many weak features that make the spectrum very crowded, with a density of about 45 lines/ $cm^{-1}$ . In fact, some of the weaker absorptions are due to the hot bands originating from the  $\nu_4 = 1$  state that at room temperature has a population of ca. 13% with respect to the ground state. The second aspect is that the Voigt profile model well reproduces both the self-broadened and the foreign broadened spectra within the experimental conditions, as expected on the basis of previous investigations. Under this point of view, it should be noted that the experimental conditions adopted coupled to the high density of lines make the use of refined line shape models, accounting for narrowing effects, to give no improvements on the retrieved parameters, also in view of their computational cost, an aspect already pointed out in previous investigations [31–33]. During the fits of the recorded spectra, resonant frequencies ( $\nu_0$ ), Lorentzian (i.e. collisional) half width ( $\Gamma_L$ ) and line intensities ( $S$ ) have been refined. Self-broadening experiments have led to the determination of self-broadening coefficients ( $\gamma_{self}^0$ ) and integrated absorption coefficients ( $S^0$ ) by a weighted linear fit (according to the reciprocal of the squared errors derived from the line shape analysis) of the collisional half widths and line intensities against the pressure of  $P_{CH_2F_2}$ , respectively:

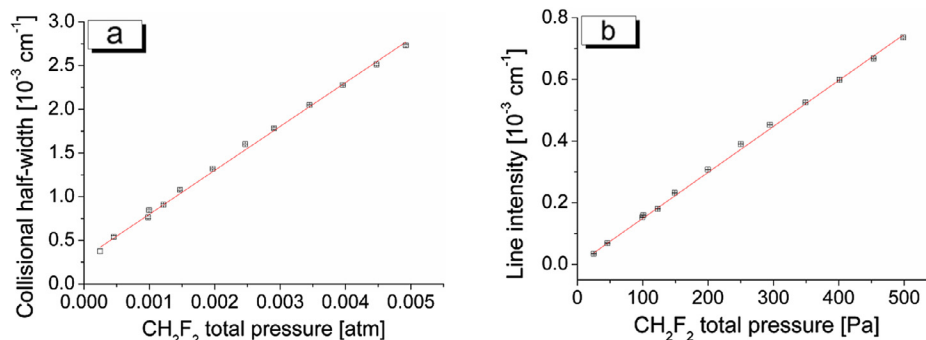
$$\Gamma_L = \Gamma_0 + \gamma_{self}^0 \times P_{CH_2F_2} \quad (2)$$

$$S = S^0 \times P_{CH_2F_2} \quad (3)$$

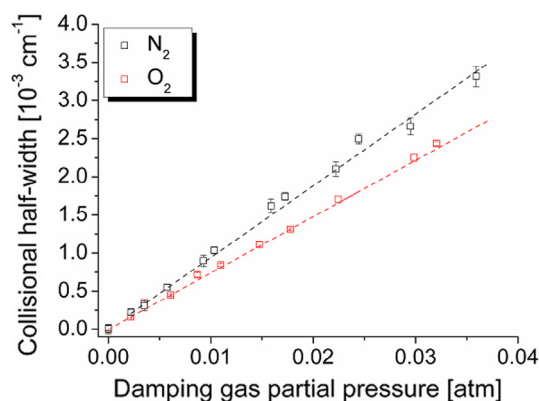
where  $\Gamma_0$  takes into account the broadening due to the Lorentzian instrumental line shape function. Examples concerning the pressure dependence of the collisional half width and the line intensity of the  $20_{11,9} \leftarrow 19_{10,9}$  ro-vibrational transition centered at 1221.6437  $cm^{-1}$ , and the corresponding linear fits, are given in Fig. 3a and b, respectively.

Foreign broadening experiments have led to the determination of  $CH_2F_2$ - $N_2$  ( $\gamma_{N_2}^0$ ) and  $CH_2F_2$ - $O_2$  ( $\gamma_{O_2}^0$ ) broadening coefficients, which have been obtained by a weighted linear fit as in Eq. (2), with the only difference that, in this case,  $\Gamma_0$  accounts for both the instrumental broadening and the self-broadening contributions. The linear fits carried out for retrieving the foreign broadening parameters of the  $20_{11,9} \leftarrow 19_{10,9}$  transition of the  $\nu_7$  band are reproduced in Fig. 4.

The line-by-line parameters obtained for the  $\nu_7$  band transitions of  $CH_2F_2$  analyzed in the present work are collected in Table 1, where the  $CH_2F_2$ -air broadening coefficients are also reported.



**Fig. 3.** (a) Self-broadening of the  $20_{11,9} \leftarrow 19_{10,9}$  ro-vibrational transition belonging to the  $\nu_7$  band of  $CH_2F_2$ , centered at 1221.6437  $cm^{-1}$  (squares: experimental half widths; straight line: linear fit of Lorentzian half widths against the  $CH_2F_2$  total pressure). Error bars are shown and are three times the uncertainties retrieved from line profile analysis. (b) Linear fit leading to the integrated absorption coefficient of the  $20_{11,9} \leftarrow 19_{10,9}$  ro-vibrational transition of  $CH_2F_2$  (squares: experimental line intensities; straight line: linear fit of line intensities against the  $CH_2F_2$  total pressure). Error bars are shown and are three times the uncertainties retrieved from line profile analysis.



**Fig. 4.** Linear fits for retrieving  $N_2$ - and  $O_2$ -broadening coefficients for the  $20_{11,9} \leftarrow 19_{10,9}$  ro-vibrational transition of the  $\nu_7$  band (squares: experimental half widths; dotted straight lines: linear fit of Lorentzian half widths against the damping gas partial pressure). The contribution due to self- and instrumental broadening has been subtracted.

These have been obtained from the measured  $N_2$ - and  $O_2$ - broadening coefficients according to the usual equation

$$\gamma_{air}^0 = 0.79\gamma_{N_2}^0 + 0.21\gamma_{O_2}^0. \quad (4)$$

From Table 1 it can be noted that  $CH_2F_2$  self-broadening coefficients range from 0.30  $cm^{-1} atm^{-1}$  to 0.66  $cm^{-1} atm^{-1}$  and take an average value of  $0.51 \pm 0.09 cm^{-1} atm^{-1}$ , thus confirming our previous findings based on a smaller subset of data [25]. In passing, it should be noted that, as in the case of other asymmetric top molecules [10,34], collisional coupling may affect the measured line shapes, particularly when even and odd doublets are degenerate, thus reducing the broadening of the spectral lines and hence leading to smaller broadening coefficients. This phenomenon arises because the unresolved  $K$  doublets are composed by two lines, namely  $J'_{K'_a K'_c} \leftarrow J''_{K''_a K''_c}$  and  $J'_{K'_a K'_c+1} \leftarrow J''_{K''_a K''_c+1}$ , which are separated by a frequency smaller than the collisional line width. When this condition is met, collisions may cause coupling between pairs of optical transitions, and population is exchanged between their initial and final states. Concerning  $N_2$ - and  $O_2$ - broadening coefficients, they vary from 0.061  $cm^{-1} atm^{-1}$  to 0.127  $cm^{-1} atm^{-1}$  and from 0.045  $cm^{-1} atm^{-1}$  to 0.085  $cm^{-1} atm^{-1}$ , respectively.

For twenty-six transitions, it is possible to compare the present results with those determined in a previous work [25]. Concerning transition frequencies, we note an overall good agreement between the two sets of data, with deviations lower or comparable



**Table 1**Assignment, transition frequency, integrated absorption coefficient, self-, nitrogen-, oxygen- and air-broadening coefficients for CH<sub>2</sub>F<sub>2</sub> transitions belonging to the  $\nu_7$  band.<sup>a</sup>

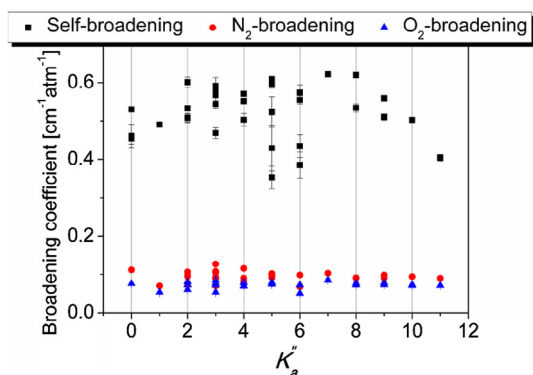
$J'$	$K'_a$	$K'_c$	$J''$	$K''_a$	$K''_c$	$\nu_0$ [cm <sup>-1</sup> ]	$S^0$ [10 <sup>-22</sup> cm molecule <sup>-1</sup> ]	$\gamma_{\text{self}}$ [cm <sup>-1</sup> atm <sup>-1</sup> ]	$\gamma_{\text{N}_2}$ [cm <sup>-1</sup> atm <sup>-1</sup> ]	$\gamma_{\text{O}_2}$ [cm <sup>-1</sup> atm <sup>-1</sup> ]	$\gamma_{\text{air}}$ [cm <sup>-1</sup> atm <sup>-1</sup> ]
18	2	17	17	1	17	1197.36528 <sub>0</sub> (6 <sub>0</sub> )	1.7 <sub>2</sub> (1 <sub>2</sub> )	0.490 <sub>7</sub> (6 <sub>3</sub> )	0.070 <sub>1</sub> (4 <sub>1</sub> )	0.053 <sub>5</sub> (2 <sub>3</sub> )	0.066 <sub>6</sub> (3 <sub>7</sub> )
39	8	31	39	7	33	1197.41165 <sub>3</sub> (4 <sub>9</sub> )	0.86 <sub>0</sub> (7 <sub>2</sub> )	0.389 <sub>7</sub> (8 <sub>6</sub> )	0.073 <sub>2</sub> (4 <sub>6</sub> )	0.045 <sub>0</sub> (3 <sub>3</sub> )	0.067 <sub>3</sub> (4 <sub>3</sub> )
23	4	20	23	1	22	1197.4496 <sub>9</sub> (2 <sub>8</sub> )	0.343 <sub>6</sub> (7 <sub>2</sub> )	0.30 <sub>2</sub> (2 <sub>6</sub> )	/	/	/
38	8	31	38	7	31	1197.47817 <sub>1</sub> (7 <sub>9</sub> )	0.85 <sub>8</sub> (6 <sub>1</sub> )	0.51 <sub>8</sub> (2 <sub>2</sub> )	0.109 <sub>5</sub> (2 <sub>6</sub> )	0.064 <sub>2</sub> (2 <sub>4</sub> )	0.100 <sub>0</sub> (2 <sub>6</sub> )
10	5	5	9	4	5	1197.52355 <sub>0</sub> (9 <sub>2</sub> ) <sup>b</sup>	5.4 <sub>6</sub> (4 <sub>4</sub> )	0.571 <sub>1</sub> (6 <sub>1</sub> )	0.089 <sub>8</sub> (8 <sub>1</sub> )	0.069 <sub>7</sub> (6 <sub>0</sub> )	0.0856(7 <sub>7</sub> )
38	8	30	38	7	32	1197.53165 <sub>3</sub> (5 <sub>5</sub> )	0.52 <sub>7</sub> (6 <sub>4</sub> )	0.56 <sub>4</sub> (3 <sub>3</sub> )	0.073 <sub>0</sub> (5 <sub>3</sub> )	0.064 <sub>0</sub> (4 <sub>3</sub> )	0.071 <sub>1</sub> (5 <sub>1</sub> )
20	12	8	21	11	10	1197.58201 <sub>0</sub> (5 <sub>7</sub> ) <sup>b</sup>	0.56 <sub>4</sub> (5 <sub>3</sub> )	0.43 <sub>2</sub> (1 <sub>7</sub> )	/	/	/
10	3	7	9	0	9	1197.59039 <sub>0</sub> (9 <sub>9</sub> )	0.12 <sub>7</sub> (5 <sub>1</sub> )	0.45 <sub>4</sub> (1 <sub>4</sub> )	/	/	/
19	1	18	18	0	18	1197.6021 <sub>1</sub> (1 <sub>3</sub> )	1.9 <sub>3</sub> (2 <sub>0</sub> )	0.530 <sub>5</sub> (7 <sub>4</sub> )	0.112 <sub>2</sub> (5 <sub>1</sub> )	0.0760 <sub>8</sub> (8 <sub>0</sub> )	0.104 <sub>6</sub> (4 <sub>2</sub> )
37	8	30	37	7	30	1197.61334 <sub>1</sub> (7 <sub>9</sub> )	0.42 <sub>7</sub> (8 <sub>7</sub> )	0.39 <sub>9</sub> (2 <sub>9</sub> )	0.071 <sub>0</sub> (6 <sub>6</sub> )	0.068 <sub>0</sub> (5 <sub>4</sub> )	0.062 <sub>4</sub> (6 <sub>4</sub> )
15	4	11	14	3	11	1197.6409 <sub>6</sub> (1 <sub>2</sub> )	4.0 <sub>3</sub> (2 <sub>7</sub> )	0.577 <sub>6</sub> (5 <sub>6</sub> )	0.108 <sub>2</sub> (2 <sub>2</sub> )	0.0779 <sub>8</sub> (5 <sub>5</sub> )	0.101 <sub>9</sub> (1 <sub>8</sub> )
37	8	29	37	7	31	1197.6515 <sub>1</sub> (1 <sub>8</sub> )	0.79 <sub>7</sub> (3 <sub>8</sub> )	0.44 <sub>8</sub> (2 <sub>5</sub> )	0.077 <sub>3</sub> (7 <sub>1</sub> )	0.071 <sub>3</sub> (6 <sub>1</sub> )	0.076 <sub>0</sub> (6 <sub>9</sub> )
43	4	39	42	5	37	1197.6940 <sub>8</sub> (3 <sub>2</sub> )	0.151 <sub>2</sub> (5 <sub>2</sub> )	0.42 <sub>9</sub> (6 <sub>0</sub> )	/	/	/
36	8	29	36	7	29	1197.74212(1 <sub>2</sub> )	0.72 <sub>6</sub> (5 <sub>6</sub> )	0.457 <sub>2</sub> (9 <sub>8</sub> )	0.081 <sub>8</sub> (4 <sub>4</sub> )	0.069 <sub>9</sub> (5 <sub>9</sub> )	0.079 <sub>3</sub> (4 <sub>7</sub> )
21	3	18	20	2	18	1197.76033 <sub>0</sub> (7 <sub>6</sub> )	4.8 <sub>2</sub> (2 <sub>8</sub> )	0.533 <sub>3</sub> (5 <sub>4</sub> )	0.106 <sub>8</sub> (5 <sub>6</sub> )	0.081 <sub>5</sub> (1 <sub>0</sub> )	0.101 <sub>5</sub> (4 <sub>6</sub> )
36	8	28	36	7	30	1197.7685 <sub>1</sub> (1 <sub>3</sub> )	0.44 <sub>4</sub> (5 <sub>0</sub> )	0.41 <sub>5</sub> (5 <sub>8</sub> )	0.091 <sub>5</sub> (8 <sub>0</sub> )	0.0590 <sub>0</sub> (6 <sub>0</sub> )	0.084 <sub>7</sub> (6 <sub>4</sub> )
24	4	21	24	1	23	1197.7997 <sub>6</sub> (2 <sub>9</sub> )	0.221 <sub>2</sub> (4 <sub>0</sub> )	0.41 <sub>5</sub> (2 <sub>8</sub> )	/	/	/
24	13	11	25	12	13	1197.82663 <sub>5</sub> (7 <sub>1</sub> ) <sup>b</sup>	0.26 <sub>3</sub> (2 <sub>6</sub> )	0.64 <sub>5</sub> (4 <sub>2</sub> )	/	/	/
15	4	12	14	3	12	1197.8357 <sub>7</sub> (1 <sub>2</sub> )	2.5 <sub>1</sub> (1 <sub>2</sub> )	0.575 <sub>5</sub> (9 <sub>3</sub> )	0.104 <sub>5</sub> (6 <sub>1</sub> )	0.072 <sub>0</sub> (1 <sub>6</sub> )	0.097 <sub>7</sub> (5 <sub>1</sub> )
6	6	0	5	5	0	1197.86850 <sub>3</sub> (9 <sub>0</sub> ) <sup>b,c</sup>	7.3 <sub>0</sub> (1 <sub>2</sub> )	0.609 <sub>1</sub> (4 <sub>7</sub> )	0.098 <sub>2</sub> (2 <sub>4</sub> )	0.078 <sub>4</sub> (2 <sub>4</sub> )	0.094 <sub>1</sub> (2 <sub>4</sub> )
35	8	27	35	7	29	1197.88269 <sub>1</sub> (1 <sub>4</sub> ) <sup>d</sup>	4.3 <sub>6</sub> (4 <sub>2</sub> )	0.51 <sub>1</sub> (1 <sub>6</sub> )	0.111 <sub>5</sub> (1 <sub>4</sub> )	0.061 <sub>3</sub> (1 <sub>7</sub> )	0.101 <sub>0</sub> (1 <sub>5</sub> )
18	3	16	17	2	16	1197.88274 <sub>7</sub> (3 <sub>8</sub> ) <sup>e</sup>	4.3 <sub>5</sub> (3 <sub>4</sub> )	0.50 <sub>9</sub> (1 <sub>2</sub> )	0.095 <sub>3</sub> (8 <sub>1</sub> )	0.073 <sub>8</sub> (1 <sub>1</sub> )	0.090 <sub>8</sub> (6 <sub>6</sub> )
21	2	19	21	1	19	1197.98169 <sub>9</sub> (7 <sub>2</sub> ) <sup>f</sup>	4.9 <sub>9</sub> (3 <sub>1</sub> )	0.556 <sub>7</sub> (5 <sub>2</sub> )	0.104 <sub>3</sub> (7 <sub>0</sub> )	0.0760 <sub>3</sub> (3 <sub>7</sub> )	0.0971 <sub>6</sub> (6 <sub>3</sub> )
34	8	26	34	7	28	1197.99397 <sub>0</sub> (6 <sub>1</sub> )	0.59 <sub>9</sub> (4 <sub>4</sub> )	0.40 <sub>6</sub> (1 <sub>9</sub> )	0.087 <sub>1</sub> (3 <sub>9</sub> )	0.074 <sub>0</sub> (1 <sub>1</sub> )	0.084 <sub>4</sub> (3 <sub>3</sub> )
15	11	4	16	10	6	1198.07018 <sub>0</sub> (7 <sub>3</sub> ) <sup>b</sup>	0.20 <sub>2</sub> (2 <sub>0</sub> )	0.41 <sub>3</sub> (2 <sub>7</sub> )	/	/	/
33	8	26	33	7	26	1198.09423 <sub>3</sub> (1 <sub>5</sub> )	0.51 <sub>9</sub> (4 <sub>6</sub> )	0.49 <sub>4</sub> (1 <sub>5</sub> )	0.087 <sub>2</sub> (2 <sub>9</sub> )	0.070 <sub>1</sub> (4 <sub>6</sub> )	0.083 <sub>6</sub> (3 <sub>3</sub> )
33	8	25	33	7	27	1198.1022 <sub>8</sub> (1 <sub>1</sub> )	1.27 <sub>0</sub> (5 <sub>9</sub> )	0.545 <sub>1</sub> (7 <sub>4</sub> )	0.090 <sub>2</sub> (3 <sub>7</sub> )	0.072 <sub>2</sub> (1 <sub>2</sub> )	0.086 <sub>4</sub> (3 <sub>2</sub> )
11	5	6	10	4	6	1198.15214 <sub>0</sub> (7 <sub>0</sub> ) <sup>b</sup>	7.0 <sub>7</sub> (2 <sub>2</sub> )	0.570 <sub>7</sub> (6 <sub>4</sub> )	0.115 <sub>9</sub> (5 <sub>3</sub> )	0.0789 <sub>1</sub> (9 <sub>2</sub> )	0.108 <sub>1</sub> (4 <sub>4</sub> )
16	4	12	15	3	12	1198.19049 <sub>7</sub> (1 <sub>5</sub> )	2.9 <sub>0</sub> (1 <sub>2</sub> )	0.56 <sub>7</sub> (1 <sub>6</sub> )	0.127 <sub>0</sub> (2 <sub>1</sub> )	0.085 <sub>0</sub> (1 <sub>0</sub> )	0.118 <sub>2</sub> (1 <sub>9</sub> )
32	8	25	32	7	25	1198.20224 <sub>7</sub> (8 <sub>1</sub> )	1.7 <sub>7</sub> (2 <sub>3</sub> )	0.557 <sub>6</sub> (6 <sub>2</sub> )	0.097 <sub>2</sub> (9 <sub>0</sub> )	0.0804 <sub>6</sub> (8 <sub>5</sub> )	0.0936 <sub>8</sub> (7 <sub>3</sub> )
32	8	24	32	7	26	1198.20769 <sub>0</sub> (2 <sub>7</sub> )	0.90 <sub>3</sub> (6 <sub>9</sub> )	0.66 <sub>1</sub> (3 <sub>1</sub> )	0.102 <sub>6</sub> (9 <sub>7</sub> )	0.079 <sub>2</sub> (2 <sub>2</sub> )	0.097 <sub>7</sub> (8 <sub>1</sub> )
11	9	2	11	8	4	1202.69706 <sub>1</sub> (1 <sub>1</sub> ) <sup>b</sup>	1.7 <sub>4</sub> (2 <sub>4</sub> )	0.49 <sub>7</sub> (3 <sub>1</sub> )	0.083 <sub>9</sub> (1 <sub>4</sub> )	0.074 <sub>7</sub> (4 <sub>1</sub> )	0.082 <sub>0</sub> (2 <sub>0</sub> )
9	7	2	8	6	2	1202.72343 <sub>9</sub> (7 <sub>4</sub> ) <sup>b</sup>	6.8 <sub>4</sub> (4 <sub>4</sub> )	0.554 <sub>9</sub> (9 <sub>9</sub> )	0.098 <sub>5</sub> (3 <sub>9</sub> )	0.072 <sub>6</sub> (2 <sub>3</sub> )	0.093 <sub>1</sub> (3 <sub>6</sub> )
10	9	1	10	8	3	1202.73245 <sub>0</sub> (6 <sub>7</sub> ) <sup>b</sup>	1.6 <sub>3</sub> (1 <sub>6</sub> )	0.60 <sub>0</sub> (4 <sub>9</sub> )	0.100 <sub>3</sub> (5 <sub>3</sub> )	0.077 <sub>1</sub> (5 <sub>4</sub> )	0.0954 <sub>3</sub> (5 <sub>3</sub> )
27	3	24	26	2	24	1202.74265 <sub>6</sub> (1 <sub>8</sub> )	2.71 <sub>8</sub> (7 <sub>9</sub> )	0.60 <sub>1</sub> (1 <sub>4</sub> )	0.098 <sub>2</sub> (5 <sub>0</sub> )	0.072 <sub>6</sub> (4 <sub>2</sub> )	0.092 <sub>8</sub> (4 <sub>2</sub> )
14	6	8	13	5	8	1202.91639 <sub>9</sub> (3 <sub>6</sub> ) <sup>b</sup>	6.2 <sub>8</sub> (4 <sub>4</sub> )	0.595 <sub>8</sub> (9 <sub>3</sub> )	0.102 <sub>2</sub> (6 <sub>7</sub> )	0.074 <sub>3</sub> (2 <sub>6</sub> )	0.096 <sub>3</sub> (5 <sub>8</sub> )
24	1	23	23	0	23	1202.93081 <sub>1</sub> (1 <sub>1</sub> )	0.59 <sub>4</sub> (1 <sub>3</sub> )	0.46 <sub>0</sub> (3 <sub>1</sub> )	/	/	/
26	4	22	25	3	22	1202.95887 <sub>2</sub> (4 <sub>3</sub> )	1.80 <sub>8</sub> (8 <sub>8</sub> )	0.46 <sub>9</sub> (1 <sub>4</sub> )	0.070 <sub>9</sub> (5 <sub>3</sub> )	0.053 <sub>6</sub> (3 <sub>9</sub> )	0.067 <sub>3</sub> (5 <sub>0</sub> )
24	3	22	23	2	22	1202.98587 <sub>0</sub> (7 <sub>8</sub> )	1.8 <sub>6</sub> (1 <sub>5</sub> )	0.50 <sub>7</sub> (1 <sub>2</sub> )	0.076 <sub>6</sub> (6 <sub>9</sub> )	0.060 <sub>7</sub> (2 <sub>6</sub> )	0.073 <sub>2</sub> (6 <sub>0</sub> )
19	5	14	18	4	14	1202.99997 <sub>0</sub> (3 <sub>8</sub> )	3.36 <sub>0</sub> (7 <sub>5</sub> )	0.551 <sub>7</sub> (6 <sub>6</sub> )	0.081 <sub>5</sub> (2 <sub>6</sub> )	0.069 <sub>7</sub> (2 <sub>1</sub> )	0.079 <sub>0</sub> (2 <sub>5</sub> )
19	5	15	18	4	15	1203.06831 <sub>5</sub> (7 <sub>6</sub> )	2.12 <sub>2</sub> (9 <sub>1</sub> )	0.50 <sub>3</sub> (1 <sub>7</sub> )	0.079 <sub>9</sub> (4 <sub>2</sub> )	0.070 <sub>6</sub> (1 <sub>6</sub> )	0.077 <sub>9</sub> (3 <sub>7</sub> )
23	4	20	22	3	20	1203.16559 <sub>9</sub> (3 <sub>5</sub> )	3.14 <sub>4</sub> (3 <sub>7</sub> )	0.54 <sub>3</sub> (1 <sub>1</sub> )	0.089 <sub>1</sub> (1 <sub>0</sub> )	0.071 <sub>8</sub> (1 <sub>8</sub> )	0.085 <sub>5</sub> (1 <sub>2</sub> )
25	15	10	26	14	12	1203.23767 <sub>9</sub> (9 <sub>0</sub> ) <sup>b</sup>	0.31 <sub>2</sub> (3 <sub>9</sub> )	0.58 <sub>4</sub> (1 <sub>8</sub> )	0.108 <sub>8</sub> (8 <sub>2</sub> )	0.062 <sub>0</sub> (4 <sub>3</sub> )	0.099 <sub>0</sub> (7 <sub>4</sub> )
44	10	35	44	9	35	1203.25383 <sub>3</sub> (2 <sub>7</sub> ) <sup>b</sup>	0.251 <sub>0</sub> (5 <sub>6</sub> )	0.33 <sub>2</sub> (1 <sub>4</sub> )	/	/	/
29	16	13	30	15	15	1203.3021 <sub>1</sub> (4 <sub>5</sub> ) <sup>b</sup>	0.249 <sub>3</sub> (8 <sub>2</sub> )	0.37 <sub>3</sub> (2 <sub>5</sub> )	/	/	/
10	7	3	9	6	3	1203.3554 <sub>5</sub> (1 <sub>1</sub> ) <sup>b</sup>	7.11 <sub>6</sub> (6 <sub>9</sub> )	0.571 <sub>8</sub> (8 <sub>1</sub> )	0.100 <sub>5</sub> (1 <sub>8</sub> )	0.0733 <sub>8</sub> (8 <sub>2</sub> )	0.094 <sub>8</sub> (1 <sub>6</sub> )
35	4	32	34	3	32	1212.9828 <sub>1</sub> (1 <sub>3</sub> )	0.43 <sub>1</sub> (4 <sub>7</sub> )	0.59 <sub>2</sub> (2 <sub>2</sub> )	0.092 <sub>9</sub> (6 <sub>3</sub> )	0.072 <sub>8</sub> (3 <sub>9</sub> )	0.088 <sub>7</sub> (5 <sub>8</sub> )
31	6	26	30	5	26	1212.98804 <sub>1</sub> (5 <sub>2</sub> )	0.756 <sub>4</sub> (5 <sub>7</sub> )	0.35 <sub>3</sub> (3 <sub>0</sub> )	/	/	/
26	7	19	25	6	19	1213.04784 <sub>1</sub> (1 <sub>9</sub> ) <sup>b</sup>	0.97 <sub>0</sub> (5 <sub>8</sub> )	0.57 <sub>4</sub> (2 <sub>0</sub> )	0.068 <sub>3</sub> (6 <sub>4</sub> )	0.0508 <sub>1</sub> (7 <sub>3</sub> )	0.064 <sub>6</sub> (5 <sub>2</sub> )
11	10	1	10	9	1	1213.05226 <sub>0</sub> (3 <sub>1</sub> ) <sup>b</sup>	6.0 <sub>5</sub> (1 <sub>1</sub> )	0.510 <sub>4</sub> (8 <sub>5</sub> )	0.098 <sub>7</sub> (4 <sub>3</sub> )	0.0780 <sub>3</sub> (5 <sub>0</sub> )	0.094 <sub>3</sub> (3 <sub>5</sub> )
16	9	7	15	8	7	1213.08295 <sub>6</sub> (4 <sub>0</sub> ) <sup>b</sup>	3.2 <sub>5</sub> (5 <sub>0</sub> )	0.53 <sub>4</sub> (1 <sub>1</sub> )	0.090 <sub>7</sub> (2 <sub>8</sub> )	0.077 <sub>2</sub> (2 <sub>7</sub> )	0.087 <sub>9</sub> (2 <sub>8</sub> )
21	8	13	20	7	13	1213.09044 <sub>3</sub> (5 <sub>0</sub> ) <sup>b</sup>	4.24 <sub>0</sub> (8 <sub>5</sub> )	0.621 <sub>8</sub> (7 <sub>1</sub> )	0.103 <sub>4</sub> (2 <sub>6</sub> )	0.085 <sub>2</sub> (4 <sub>8</sub> )	0.099 <sub>6</sub> (3 <sub>0</sub> )
32	6	26	31	5	26	1213.09904 <sub>5</sub> (6 <sub>1</sub> )	0.43 <sub>5</sub> (6 <sub>1</sub> )	0.52 <sub>3</sub> (3 <sub>9</sub> )	0.092 <sub>2</sub> (4 <sub>7</sub> )	0.079 <sub>1</sub> (7 <sub>7</sub> )	0.089 <sub>4</sub> (5 <sub>4</sub> )
16	15	2	16	14	2	1221.32998 <sub>3</sub> (3 <sub>1</sub> ) <sup>b</sup>	0.203 <sub>6</sub> (4 <sub>1</sub> )	0.32 <sub>0</sub> (2 <sub>2</sub> )	/	/	/
30	9	21	29	8	21	1221.3891 <sub>6</sub> (2 <sub>8</sub> ) <sup>b</sup>	1.4 <sub>9</sub> (1 <sub>1</sub> )	0.619 <sub>7</sub> (7 <sub>6</sub> )	0.090 <sub>9</sub> (8 <sub>4</sub> )	0.072 <sub>3</sub> (4 <sub>2</sub> )	0.087 <sub>0</sub> (7 <sub>5</sub> )
41	7	35	40	6	35	1221.5257 <sub>0</sub> (2 <sub>6</sub> )	0.185 <sub>1</sub> (4 <sub>6</sub> )	0.38 <sub>5</sub> (3 <sub>5</sub> )	/	/	/
25	10	15	24	9	15	1221.53119 <sub>3</sub> (5 <sub>1</sub> ) <sup>b,g</sup>	1.71 <sub>1</sub> (9 <sub>3</sub> )	0.559 <sub>3</sub> (7 <sub>5</sub> )	0.090 <sub>8</sub> (7 <sub>2</sub> )	0.073 <sub>8</sub> (2 <sub>1</sub> )	0.087 <sub>1</sub> (6 <sub>1</sub> )
43	7	36	42	6	36	1221.6149 <sub>5</sub> (4 <sub>8</sub> )	0.189 <sub>4</sub> (5 <sub>0</sub> )	0.43 <sub>5</sub> (3 <sub>0</sub> )	/	/	/
20	11	9	19	10	9	1221.6437 <sub>6</sub> (8 <sub>0</sub> ) <sup>b</sup>	2.05 <sub>6</sub> (2 <sub>8</sub> )	0.502 <sub>5</sub> (6 <sub>8</sub> )	0.094 <sub>0</sub> (2 <sub>1</sub> )	0.073 <sub>4</sub> (1 <sub>2</sub> )	0.089 <sub>7</sub> (1 <sub>9</sub> )
24	12	12	23	11	12	1227.13560 <sub>1</sub> (2 <sub>2</sub> ) <sup>b</sup>	1.00 <sub>0</sub> (3 <sub>8</sub> )	0.404 <sub>7</sub> (8 <sub>6</sub> )	0.090 <sub>3</sub> (5 <sub>5</sub> )	0.071 <sub>5</sub> (2 <sub>7</sub> )	0.086 <sub>3</sub> (4 <sub>9</sub> )

<sup>a</sup> Figures in parentheses are statistical errors in the units of the last significant digits.<sup>b</sup> Even and odd components of the transition.<sup>c</sup> Overlapped with the 35<sub>8,28</sub> ← 35<sub>7,28</sub> transition.<sup>d</sup> Overlapped with the 18<sub>3,16</sub> ← 17<sub>2,16</sub> transition.<sup>e</sup> Overlapped with the 35<sub>8,27</sub> ← 35<sub>7,29</sub> transition.<sup>f</sup> Overlapped with the 34<sub>8,27</sub> ← 34<sub>7,27</sub> transition.<sup>g</sup> Overlapped with the 41<sub>7,35</sub> ← 40<sub>6,35</sub> transition.

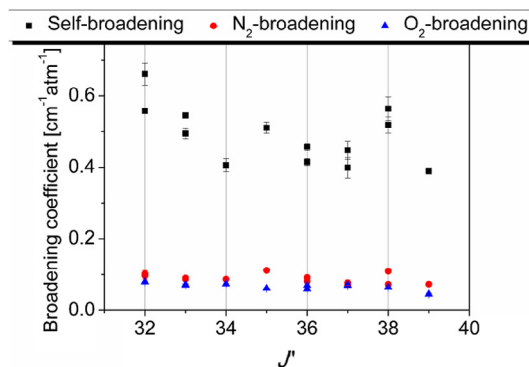
to the accuracy of the calibration. Exceptions occur for eight transitions (10<sub>5,5</sub> ← 9<sub>4,5</sub>; 38<sub>8,30</sub> ← 38<sub>7,32</sub>; 20<sub>12,8</sub> ← 21<sub>11,10</sub>; 19<sub>1,8</sub> ← 18<sub>0,8</sub>; 37<sub>8,30</sub> ← 37<sub>7,30</sub>; 15<sub>4,11</sub> ← 14<sub>3,11</sub>; 37<sub>8,29</sub> ← 37<sub>7,31</sub>; 21<sub>3,18</sub> ← 20<sub>2,18</sub>) for which the differences are larger than  $1.5 \times 10^{-3}$  cm<sup>-1</sup>. This is likely due to

of view, the agreement with previous measurements is good, around 10%, and it reflects the statistical errors affecting the intensities. In passing, it should be noted that in Ref. [25] there is a typo in the value reported for the  $37_{8,30} \leftarrow 37_{7,30}$  transition intensity ( $0.253 \times 10^{-22}$  cm molecule $^{-1}$ ), the correct one being  $0.52_3(1_3) \times 10^{-22}$  cm molecule $^{-1}$ . Concerning self-broadening coefficients, the agreement between the two datasets is good, on average within 6% and in general within 10%. The largest discrepancies of 21% and 25% are for the transitions  $38_{8,30} \leftarrow 38_{7,32}$  and  $36_{8,28} \leftarrow 36_{7,30}$ , respectively. Yet, the values here determined provide a smoother variation of the self-broadening coefficients with the  $J''$  quantum number within the  $K''_a = 7$  subbranch, and hence they are expected to be more reliable. The situation for  $N_2$ - and  $O_2$ -broadening coefficients is more involved and there is a larger number of data that presents deviations above 15%, particularly in the case of collisions between  $CH_2F_2$  and  $O_2$ . This is not surprising considering that broadening parameters are strongly affected by the location of the baseline. In particular, it has been pointed out that an error in the base of 1% may cause an error of 12% in the retrieved Lorentzian line-width [35]. Due to the high density of transitions in the spectra analyzed and the line overlaps induced by the pressurization with the buffer gas, the location of the base line can be expected to be the main source of uncertainty in the retrieved foreign broadening coefficients, an aspect that is only partially accounted for in the errors reported in Table 1, which are of statistical nature. While the  $N_2$ - and  $O_2$ - broadening coefficients here determined appear consistent in view of the smoother variation with respect to the rotational quantum numbers, a conservative estimate of their uncertainty is 10–15%, which, taken into account the difficulty in the analysis caused by the spectral congestion, appears more than satisfactory.

From the collisional cross sections obtained experimentally, it has been possible to obtain some hints about the  $K_a$  dependence of R-branch transitions and the  $J$  dependence for Q-type transitions belonging to the  $K''_a = 7$  sub-branch. The observed trends are illustrated in Figs. 5 and 6, respectively. In the case of the self-broadening of R-branch transitions, Fig. 5 suggests an initial increase of the broadening efficiency with the  $K_a$  quantum number, up to  $K''_a \approx 5$ , and then a decrease as  $K''_a$  raises further. As expected, the  $K_a$  dependence of  $N_2$ - and  $O_2$ -broadening parameters is much weaker, and within the  $K_a$ -range here investigated, they keep an almost constant value around  $0.094 \pm 0.012$  and  $0.0718 \pm 0.0095$  cm $^{-1}$  atm $^{-1}$ , respectively. The  $J$ -dependence of self-broadening coefficients for Q-branch transitions belonging to the  $K''_a = 7$  subbands is more involved: from Fig. 6, it can be speculated that they have the tendency to decrease with increasing  $J''$ , at least within the range  $32 \leq J'' \leq 39$ . Concerning the  $CH_2F_2$ - $N_2$  and  $CH_2F_2$ - $O_2$  collisional cross sections, the dependence on  $J$  is very weak,



**Fig. 5.** Dependence of self-,  $N_2$ - and  $O_2$ - broadening coefficients on the  $K_a$  pseudo-quantum number for  $CH_2F_2$  ro-vibrational transitions belonging to the R-branch of the  $\nu_7$  band.



**Fig. 6.** Dependence of self-,  $N_2$ - and  $O_2$ - broadening coefficients on the  $J$  quantum number for  $CH_2F_2$  Q-branch ro-vibrational transitions belonging to the  $K''_a = 7$  manifold of the  $\nu_7$  band.

although the efficiency of collisions in broadening the spectral lines seems to decrease as  $J''$  increases. However, it should be pointed out that the trend suggested for this dependence has to be considered as indicative, since the  $J$ -range accessed by the present measurements is quite small, to allow definitive conclusion.

#### 4. Conclusion

Spectroscopic parameters are crucial elements for remote sounding techniques and for understanding of the Earth's atmospheric properties, also in relation to global climate changes caused by anthropogenic activities. Under this point of view, HFCs represent trace gas pollutants that are able to alter the Earth's radiative balance thus contributing to global warming. In the present work, the line shape parameters of difluoromethane have been determined for several ro-vibrational transitions within the atmospheric window region around 8.2–8.3  $\mu$ m. The measurements have been carried out exploiting tunable diode laser spectroscopy by perturbing  $CH_2F_2$  by itself, as well as by the atmospherically relevant buffer gases  $N_2$  and  $O_2$ . The recorded transitions belong to the  $\nu_7$  band, which is one of the strongest bands of the  $CH_2F_2$  IR spectrum, and the study has been complicated by the high-density of lines in the high resolution spectra of this molecule. In spite of this, the line profile analysis has led to the determination, for 60 lines, of resonant frequencies, integrated absorption coefficients (i.e. line intensities) and self-,  $N_2$ - and  $O_2$ -broadening coefficients. In addition, from the measured  $N_2$ - and  $O_2$ - broadening parameters, the broadening coefficients of  $CH_2F_2$  in air have been derived. Finally, from the obtained collisional cross sections some indications about their quantum number dependence have been drawn.

As a concluding remark it is worth pointing out that, despite the complication arising from the spectral congestion, studies aiming at the determination of  $CH_2F_2$  broadening parameters should be encouraged, because this molecule actually represents an important atmospheric trace gas pollutants. For this reason it has been widely investigated by the spectroscopic community, in both the IR and MW spectral regions, with the aim of understanding the rotational/vibrational energy level patterns and deriving set of spectroscopic parameters for simulating its spectra, while very little work has been devoted to broadening parameters, with the present contribution being an effort for bridging this gap.

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